# Vacuum Polarization and Casimir Effect within (3+1)D Maxwell-Chern-Simons Electrodynamics with Lorentz violation

A. F. Bubnov, O. G. Kharlanov, and V. Ch. Zhukovsky

Moscow State University, Department of Theoretical Physics

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# Outline

### Introduction

- Casimir effect within extended QED (O.G.Kharlanov and V.Ch.Zhukovsky)
  - Vacuum energy via the zeta function regularization
  - Vacuum energy via the residue theorem
  - Conclusion
- 3 Effective action in QED under the Lorentz violation (A.F.Bubnov and V.Ch.Zhukovsky)
  - The Model
  - Induced Chern-Simons term in the constant field
  - Quadratic contribution
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# Standard Model Extension (SME [Kostelecký])

- Elaborated for studying the manifestation of the 'New Physics' (Strings, Extra Dimensions, Quantum Gravity,...) at low energies  $E \ll m_{\rm Pl} \sim 10^{19} \Gamma$  pB
- Introduces a set of correction terms to the Lagrangian of SM (no new fields!), that maintain some 'natural' features of SM:
  - observer Lorentz invariance (although the vacuum is not Lorentz-invariant)
  - unitarity
  - microcausality
  - $SU(3)_C \times SU(2)_I \times U(1)_Y$  gauge invariance
  - $\bullet\,$  power-counting renormalizabliity (for the minimal  ${\rm SME})$
- When  $E \ll m_W \sim$ , the SME results in the extended QED with  $U(1)_{em}$  gauge invariance typical for SM



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## (3+1)D Maxwell-Chern-Simons electrodynamics

A particular case of extended QED with the Chern-Simons term:

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}\eta^{\mu}\epsilon_{\mu\nu\alpha\beta}A^{\nu}F^{\alpha\beta} + \bar{\psi}(i\gamma^{\mu}D_{\mu} - m)\psi.$$

 $\eta^{\mu}$  is a constant 4-vector, breaks  ${\rm CPT}$  and Lorentz invariance.

 $\eta^{\mu}$  may be a manifestation of axion condensation [Carroll, Field, Jackiw,1992], or of the background torsion [Dobado,Maroto,1996]



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(3+1)D Maxwell-Chern-Simons electrodyanamics,  $\eta^{\mu} = \{\eta, \mathbf{0}\}$ 

- Photon sector:  $\psi, \bar{\psi} = 0$
- Two infinite parallel superconductor plates separated by D = 2a
- Gauge:  $A^0 = 0$ , div  $\mathbf{A} = 0$
- Equations of motion:  $\Box \mathbf{A} = 2\eta \operatorname{rot} \mathbf{A}$
- $T^{00} = \frac{1}{2}(\mathbf{E}^2 + \mathbf{H}^2) \eta \mathbf{A} \cdot \mathbf{H}$

When  $a < \pi/4 |\eta|$ , the theory is stable [1]!

Vacuum energy and Casimir force (per unit area):  $E_{vac} = \int \frac{d^3x}{L^2} \langle T^{00}(x) \rangle = \sum_n \frac{\omega_n(D)}{2L^2}, \quad f_{Casimir} = \frac{1}{L^2} \frac{\partial E_{vac}}{\partial D},$ *n* is a complete set of quantum numbers,  $L \to \infty$  is linear plate size.

The force is gauge-invariant, athough the energy is not [1]



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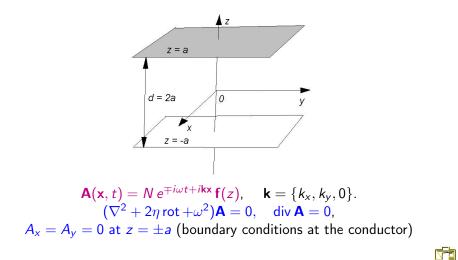
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# One-photon eigenstates [1]



# One-photon eigenstates [2]

Ansatz:

$$\begin{split} \mathbf{A}_{\epsilon,\mathbf{k},\Pi,n_z}(\mathbf{x},t) &= N \, e^{-i\epsilon\omega t + i\mathbf{k}\mathbf{x}} (f_z \mathbf{e}_z + f_k \hat{\mathbf{k}} + f_{zk}[\mathbf{e}_z \hat{\mathbf{k}}]), \quad \mathbf{k} = \{k_x,k_y,0\}. \end{split}$$
Transversality implies:  $f_k = \frac{i}{k} \partial_z f_z, \end{split}$ 

Parity  $\Pi = \pm 1$ :  $f_k(-z) = -\Pi f_k(z), f_{zk,z}(-z) = \Pi f_{zk,z}(z).$ 

Equations for  $f_{kz,z}$ :

$$\begin{aligned} (\omega^2 - k^2 + \partial_z^2)kf_y &= 2i\eta(k^2 - \partial_z^2)f_z\\ (\omega^2 - k^2 + \partial_z^2)f_z &= -2i\eta kf_y,\\ f_y(a) &= 0, \qquad \partial_z f_z(a) = 0 \end{aligned}$$

The existence of nontrivial solutions implies that:

 $g_{\Pi}(\omega^2) \equiv \varphi_{\Pi}(\varkappa_+ a) \varphi_{-\Pi}(\varkappa_- a) \sin \theta_- + \varphi_{\Pi}(\varkappa_- a) \varphi_{-\Pi}(\varkappa_+ a) \sin \theta_+ = 0,$ 

 $\varkappa_{\pm} = \sqrt{K_{\pm} - k^2}, \quad K_{\pm} = \mp \eta + \sqrt{\omega^2 + \eta^2}, \quad \sin \theta_{\pm} = \varkappa_{\pm} / K_{\pm}; \quad \varphi_{\pm 1}(x) \equiv \begin{cases} \cos x \\ \sin x \\ \sin x \\ \sin x \end{cases}$ 

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### Vacuum energy

Renormalized vacuum energy (Casimir energy) per 1cm<sup>2</sup>:

$$E_{vac} = \sum_{n} \frac{\omega_n}{2L^2} = \frac{1}{2}\zeta(-1/2), \quad f_{Casimir} = \frac{\partial E_{vac}}{\partial D},$$
  
$$\zeta(s) = \frac{1}{L^2} \sum_{n} (\omega_n^2)^{-s} = \int_0^\infty \frac{kdk}{2\pi} \sum_{\Pi = \pm 1} \sum_{n_z} (\omega_{k,\Pi,n_z}^2)^{-s}$$

For sufficiently large Re s, the series for  $\zeta(s)$  is convergent; for other  $s \in \mathbb{C}$ , it is analytically continued.

$$\zeta(s)|_{\eta=0} = \int_{0}^{\infty} \frac{kdk}{2\pi} \sum_{\Pi=\pm 1} \sum_{n_z=1}^{\infty} \left(k^2 + \left(\frac{\pi n_z}{2a}\right)^2\right)^{-s} = \left(\frac{D}{\pi}\right)^{2s-2} \frac{\zeta_R(2s-2)}{2\pi(s-1)},$$

$$f_{Casimir}|_{\eta=0} = \frac{\pi^2}{240D^4} \quad (\text{attraction}),$$

$$R(s) = \sum_{n=1}^{\infty} \frac{1}{n^s} \text{ is the Riemann zeta function.}$$
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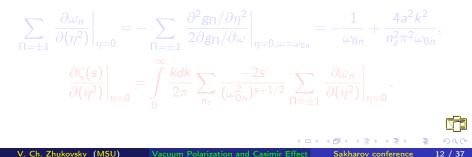
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Vacuum Polarization and Casimir Effect

### The $\eta$ -correction to the zeta function [1]

$$\zeta(s) = \int_{0}^{\infty} \frac{kdk}{2\pi} \sum_{\Pi=\pm 1} \sum_{n_z} (\omega_n^2)^{-s}, \quad n = \{k, \Pi, n_z\}$$
$$\omega_n|_{\eta=0} \equiv \omega_{0n} = \sqrt{k^2 + \left(\frac{\pi n_z}{2a}\right)^2}.$$

Note that  $\omega_n$  are the roots of the equation  $g_{\Pi}(\omega^2) = 0$ , which is even with respect to changing the sign of  $\eta$ , then  $\omega_n = \omega_n(\eta^2)$ .



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$$\sum_{\Pi=\pm 1} \frac{\partial \omega_n}{\partial (\eta^2)} \Big|_{\eta=0} = -\sum_{\Pi=\pm 1} \frac{\partial^2 g_{\Pi} / \partial \eta^2}{2 \partial g_{\Pi} / \partial \omega} \Big|_{\eta=0,\omega=\omega_{0n}} = -\frac{1}{\omega_{0n}} + \frac{4a^2k^2}{n_z^2 \pi^2 \omega_{0n}},$$
$$\frac{\partial \zeta(s)}{\partial (\eta^2)} \Big|_{\eta=0} = \int_0^\infty \frac{k dk}{2\pi} \sum_{n_z} \frac{-2s}{(\omega_{0n}^2)^{s+1/2}} \sum_{\Pi=\pm 1} \frac{\partial \omega_n}{\partial (\eta^2)} \Big|_{\eta=0}.$$

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$$\frac{1}{2\pi(s-1)} \left(\frac{D}{2s-2}\right)^{2s-2} \left(z_1(2s) + z_2(2s)\right)^{2s-2} \left(z_2(2s) + z_2(2s)\right)^{2s-2} \left$$

$$\zeta(s) = \frac{1}{2\pi(s-1)} \left(\frac{D}{\pi}\right)^{2s-2} \left(\zeta_R(2s-2) + (s-2)\left(\frac{\eta D}{\pi}\right)^2 \zeta_R(2s) + \mathcal{O}(\eta^4)\right)$$



### The correction to the Casimir force

$$f_{Casimir} = \frac{\partial}{\partial D} \frac{\zeta(-1/2)}{2} = \frac{\pi^2}{240D^4} \left( 1 + \frac{25}{3\pi^2} (\eta D)^2 + \mathcal{O}((\eta D)^4) \right), \ |\eta| D \ll 1.$$

Discussion:

- The correction is attractive, contrary to the recent result obtained by [Frank,Turan,2006]
- The difference from the Maxwell value is considerable for comparatively large  ${\cal D}$
- Experimental data [Mohideen et al.,  $D \sim 500$ nm,  $L \sim 1$ cm, 1%accuracy] gives the constraint:

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eV.

Some authors claim that sensing the Casimir force is possible at  $D \lesssim 1$ mm, then one could place a harder constraint  $|\eta| \lesssim 10^{-5}$ eV



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### Sum $\rightarrow$ complex plane integral [1]

$$f_{Casimir} = \frac{1}{L^2} \frac{\partial}{\partial D} \sum_{\omega_n \in \mathbb{R}^+} \frac{\omega_n(D)}{2} = \frac{1}{2} \frac{\partial}{\partial D} \int_0^\infty \frac{kdk}{2\pi} D(S_+ + S_-),$$

Smooth cutoff regularization:

$$S_{\Pi} = \frac{1}{D} \sum_{\omega_{k,\Pi,n_z} \in \mathbb{R}^+} \omega_{k,\Pi,n_z} e^{-\omega_{k,\Pi,n_z}/\sqrt{k\Lambda}}, \quad \Lambda \to +\infty.$$

Instead of  $g_{\Pi}(\omega^2)$  whose zeros are the one-photon energy eigenvalues, we will use the meromorphic (analytical, except for the numerable set of poles; in particular, with no branch points) function

$$\tilde{g}_{\Pi}(K_{+}) \equiv \frac{g_{\Pi}(\omega)}{\varphi_{\Pi}(\varkappa_{+}a)\varphi_{\Pi}(\varkappa_{-}a)} = \tan^{\Pi}\varkappa_{+}a\sin\theta_{+} + \tan^{\Pi}\varkappa_{-}a\sin\theta_{-},$$
$$\omega^{2} = K_{+}K_{-}, \quad K_{-} = K_{+} + 2\eta.$$

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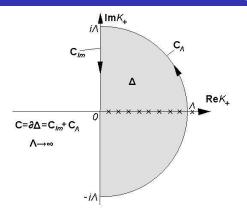
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Residue theorem (we assume  $\eta \ge 0$ , since the spectrum depends on  $|\eta|$ ):

$$\oint_{C} \frac{dK_{+}}{2\pi i} \omega \frac{\partial \tilde{g}_{\Pi} / \partial K_{+}}{\tilde{g}_{\Pi}} = S_{\Pi} D + \sum_{\bar{\omega}_{n}} \bar{\omega}_{n} \frac{\operatorname{Res}\left[\partial \tilde{g}_{\Pi} / \partial K_{+}, K_{+} = \bar{\omega}_{n}\right]}{\tilde{g}_{\Pi}(\bar{\omega}_{n})},$$
where  $\bar{\omega}_{n}$  are the poles of function  $\partial \tilde{g}_{\Pi} / \partial K_{+}$  within  $\Delta \bar{g} > 4 \bar{$ 

Vacuum Polarization and Casimir Effect

## Sum $\rightarrow$ complex plane integral [3]

Transforming the pole residue term back into an integral, we obtain:

$$\begin{split} \tilde{g}_{\Pi}(K_{+}) &\equiv \tan^{\Pi} \varkappa_{+} a \sin \theta_{+} + \tan^{\Pi} \varkappa_{-} a \sin \theta_{-}, \\ S_{\Pi} &= \frac{\Pi}{2} \oint_{C} \frac{\omega dK_{+}}{2\pi i \tilde{g}_{\Pi}(K_{+})} \left\{ 2 - \tan^{\Pi} \varkappa_{+} a \tan^{\Pi} \varkappa_{-} a \left( \frac{\sin \theta_{-}}{\sin \theta_{+}} + \frac{\sin \theta_{+}}{\sin \theta_{-}} \right) + \frac{\Pi \tan^{\Pi} \varkappa_{+} a}{\varkappa_{+} a} + \frac{\Pi \tan^{\Pi} \varkappa_{-} a}{\varkappa_{-} a} \right\} \end{split}$$

The integral over the semicircle  $C_{\Lambda}$  does not depend on a, when  $\Lambda \to \infty$ , within any finite order in a, thus it is cancelled when renormalized. Renormalization:

 $S_{\Pi}^{ren}(D) = S_{\Pi}(D) - S_{\Pi}^{div}(\infty), \quad S_{\Pi}^{div}(D) = C_1 + C_2/D \text{ at } D \to \infty.$ 



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### After renormalization and $\Lambda \to \infty$

Let us redefine  $K_+ \rightarrow -iK_+$ , and make all momentum quantities dimensionless multiplying them by *a*, then we obtain:

$$f_{Casimir} = \frac{1}{2} \frac{\partial}{\partial D} \left( D \frac{(\tilde{S}_{+} + \tilde{S}_{-})}{a^{4}} \right),$$
$$\tilde{S}_{\Pi} = -\frac{1}{2} \int_{0}^{\infty} \frac{kdk}{2\pi} \int_{\infty}^{+\infty} \frac{dK_{+}}{2\pi} \frac{\operatorname{sgn} K_{+} \sqrt{K_{+}K_{-}}}{\tanh^{\Pi} \varkappa_{+} \cosh \theta_{+} + \tanh^{\Pi} \varkappa_{-} \cosh \theta_{-}} \Sigma_{\Pi},$$

$$\begin{split} \boldsymbol{\Sigma}_{\Pi} &= 1 + \tanh^{\Pi} \varkappa_{+} \tanh^{\Pi} \varkappa_{-} \frac{\cosh \theta_{+}}{\cosh \theta_{-}} - \left(1 + \frac{\cosh \theta_{+}}{\cosh \theta_{-}}\right) \tan^{\Pi} \varkappa_{+} + \\ &+ \frac{\tanh^{\Pi} \varkappa_{+} - \tanh^{\Pi} \varkappa_{-}}{\cosh \theta_{+} + \cosh \theta_{-}} \frac{\cosh \theta_{+} \sinh^{2} \theta_{-}}{\varkappa_{-}} + \left(" - " \leftrightarrow " + "\right) \\ & K_{-} &= K_{+} - i\eta D, \quad \varkappa_{\pm} = \sqrt{k^{2} + K_{\pm}^{2}}, \quad \sinh \theta_{\pm} = k/K_{\pm}. \end{split}$$

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#### After the expansion with respect to $\eta D$ and taking the integrals, we obtain:

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## Casimir effect within extended QED (O.G.Kharlanov and V.Ch.Zhukovsky)

- Vacuum energy via the zeta function regularization
- Vacuum energy via the residue theorem
- Conclusion

# 3 Effective action in QED under the Lorentz violation (A.F.Bubnov and V.Ch.Zhukovsky)

- The Model
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## Main results

- The eigenstates and energy eigenvalues for the Maxwell-Chern-Simons photon between the conducting plates
- The vacuum is stable when  $D|\eta| < \pi/2$  [1]
- ullet The leading correction to the Casimir force, which is quadratic in  $\eta$
- Constraint on  $\eta$

References:

[1] V.Ch.Zhukovsky and O.G.Kharlanov, *Casimir effect within (3+1)D Maxwell-Chern-Simons electrodynamics*, arXiv/0905.3689[hep-th].



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### The Model

Another special case of the Extended QED; fermion sector:

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \mathcal{L}_{\psi},$$
  
 $\mathcal{L}_{\psi}[\bar{\psi}, \psi, A, b] = \bar{\psi}(i\hat{\partial} - \hat{A} + \hat{b}\gamma^{5} - m)\psi, \quad \hat{\xi} \equiv \gamma^{\mu}\xi_{\mu}.$ 

where  $b_{\mu}$  is a constant 4-vector, that violates CPT and Lorentz invariance of the theory.

Experimental constraints on  $b^{\mu}$  for the electron:

$$|b_0| \lesssim 10^{-2} \text{eV}, \quad |\mathbf{b}| \lesssim 10^{-19} \text{eV}.$$

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## Induced Chern-Simons term in the constant field [1]

 $F_{\mu\nu} = const$ 

Effective action in the proper time representation:

$$i\Gamma^{\text{eff}}[A, b] = -\frac{1}{2} \int_0^\infty \frac{dz}{z} \operatorname{Tr} e^{-zH},$$
  
$$H = -\pi^{\mu} \pi_{\mu} - 2i\sigma^{\mu\nu} b_{\mu} \pi_{\nu} \gamma^5 + \frac{1}{2} \sigma^{\mu\nu} F_{\mu\nu} + b_{\mu} b^{\mu} + m^2,$$

where  $\pi^{\mu} = i\partial_{\mu} - A_{\mu}$ ,  $\sigma^{\mu\nu} = \frac{i}{2}[\gamma^{\mu}, \gamma^{\nu}]$ ,  $\gamma^{5} = -i\gamma^{0}\gamma^{1}\gamma^{2}\gamma^{3}$ .

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In this expression, we apply the Baker-Hausdorf formula to the exponent to find its expansion into a series with respect to  $b_{\mu}$ :

 $\exp\left(\tau(A+B)\right) = \exp\left(\tau A\right) \cdot \exp\left(\tau B\right) \cdot L^{-1}(\tau), \ \frac{d\ln L}{d\tau} = B - e^{-\tau B} f(\tau) e^{\tau B}$  $A \equiv z\pi^{\mu}\pi_{\mu} - rac{1}{2}z\sigma^{\mu\nu}F_{\mu\nu}, \ B \equiv 2iz\sigma^{\mu\nu}b_{\mu}\pi_{\nu}\gamma^{5}, \ f(\tau) = e^{-\tau A}Be^{\tau A}.$ 

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$$i\Gamma^{\text{eff}}(A,b) = -\frac{1}{2} \int_0^\infty \frac{dz}{z} \operatorname{Tr} e^{-zH} = b_\alpha R^{\mu\alpha} \int_0^\infty \frac{dz}{z} \langle x | \pi_\mu \exp\left(z(\pi_\nu \pi^\nu)\right) | x \rangle$$
  
where  $R^{\mu\alpha}$  is the combination of the field tensor  $F^{\mu\nu}$  and the metric  $g^{\mu\nu}$ .

$$\int d^4x \left\langle x \right| \pi_{\mu} e^{\left(z(\pi_{\nu}\pi^{\nu})\right)} \left| x \right\rangle = P^{\rho}_{\mu} \frac{\partial}{\partial \lambda^{\rho}} \int d^4x \left\langle x \right| e^{-\frac{1}{4}z\lambda^2} e^{\left(z(\pi^{\nu}+\lambda^{\nu})^2\right)} \left| x \right\rangle \Big|_{\lambda=0}$$

vanishes due to the gauge invariance of the theory,

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Within the linear approximation, taking the trace tr, we obtain:

$$i\Gamma^{\text{eff}}(A,b) = -\frac{1}{2} \int_0^\infty \frac{dz}{z} \operatorname{Tr} e^{-zH} = \frac{b_\alpha}{R^{\mu\alpha}} \int_0^\infty \frac{dz}{z} \langle x | \pi_\mu \exp\left(z(\pi_\nu \pi^\nu)\right) | x \rangle$$

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Vacuum Polarization and Casimir Effect

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where  $R^{\mu\alpha}$  is the combination of the field tensor  $F^{\mu\nu}$  and the metric  $g^{\mu\nu}$ .

The matrix element  $\langle x | ... | x \rangle$  can be transformed into the form:

$$\int d^4x \, \langle x | \, \pi_\mu e^{(z(\pi_\nu \pi^\nu))} \, | x \rangle = P^\rho_\mu \frac{\partial}{\partial \lambda^\rho} \int d^4x \, \langle x | \, e^{-\frac{1}{4}z\lambda^2} e^{(z(\pi^\nu + \lambda^\nu)^2)} \, | x \rangle \, \Big|_{\lambda=0},$$

where  $P^{\rho}_{\mu}$  is a polynomial in the field strength. This latter expression vanishes due to the gauge invariance of the theory.

V. Ch. Zhukovsky (MSU)

Vacuum Polarization and Casimir Effect

## Induced Chern-Simons term in the constant field [3]

In the following papers, the Chern-Simons-like contribution of the form  $\beta \tilde{F}^{\mu\nu} A_{\mu} b_{\nu}$  to the effective action of QED was calculated:

- R. Jackiw and V.A. Kostelecky, *Phys. Rev. Lett* 82, 3572 (1999). The earliest publication, the coefficient  $\beta = \frac{3}{16\pi^2}$
- M.B. Hott, J.L.Tomazelli, Induced Lorentz and PCT symmetry Breaking in External Electromagnetic Field, arXiv/hep-th/9912251.  $\beta \neq 0$ , though depends on the regularization scheme.
- Y.A. Sitenko, K.Y. Rulik On the effective lagrangian in spinor electrodynamics with added violation of Lorentz and CPT symmetries, arXiv/hep-th/0212007.
- ...and others

Our calculations show that no Chern-Simons term (linear in  $b_{\mu}$ ) is induced by a fermion loop in the framework of the extended QED



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### Quadratic contribution: Magnetic field [1]

Constant homogeneous magnetic field:  $F^{12} = -F^{21} = -H$ ,  $b_{\mu} = \{b_0, 0, 0, 0\}$ .

Then the exponent  $e^{\tau(A+B)} = e^{\tau A}e^{\tau B}$  is "decoupled" in the expression for the effective action, since

$$[A,B] = 4z^2 \gamma^5 b^{\mu} \Pi_{\nu} \sigma^{\nu \alpha} F_{\alpha \mu} = 0.$$

The  $b_0^2$ -contribution to the effective action reads:

$$i\Gamma_{b_0^2}^{\text{eff}}(H,b) = -2\int_0^\infty \frac{dz}{z} \times \left(A_0 + \frac{A_1}{z}\frac{\partial}{\partial\alpha}\right) \int d^4x \left\langle x\right| e^{-z((\pi_4)^2 + \alpha(\pi_\perp^2 + \pi_\parallel^2))} \left|x\right\rangle\Big|_{\alpha=1},$$

where  $A_0,\;A_1$  are certain field combinations. The matrix element in the integrand expression

$$\int d^4x \langle x | e^{-z(\pi)_E^2} | x \rangle = \frac{H}{16\pi^2 z \sqrt{\alpha}} \frac{1}{\operatorname{sh}(\alpha z H)}.$$

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Finally we find the  $b_0^2$ -contribution to the effective action:

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and, after the renormalization,

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This integral leads to the Euler psi-function, taking the magnetic field strength  ${\cal H}$  exactly into consideration

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$$i\Gamma_{ren \ b_0^2}^{\mathrm{eff}}(H,b) = -b_0^2 \frac{m^2}{4\pi^2} \Big(\psi(m^2/2H) - \log{(m^2/2H)} + \frac{H}{m^2}\Big).$$

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### Quadratic contribution: Electric field

Constant homogeneous electric field:  $F^{03} = -F^{30} = E$ ,  $b_{\mu} = \{0, b_1, 0, 0\}.$ 

Like in the previous case, the exponent of the Hamiltonian is "decoupled" since

$$[A,B] = 4z^2 \gamma^5 b^{\mu} \Pi_{\nu} \sigma^{\nu \alpha} F_{\alpha \mu} = 0.$$

The calculations are nearly analogous to the magnetic field case and give:

$$i\Gamma_{b_1^2}^{ ext{eff}}(E,b) = -rac{b_1^2}{4\pi^2} \int_0^\infty dz e^{-zm^2} \Big(rac{E^2}{\sin^2(zE)}\Big),$$

and, after the renormalization, the effective action

$$i\Gamma_{b_1^2}^{\text{eff}}(E,b) = -\frac{b_1^2}{4\pi^2} \int_0^\infty dz e^{-zm^2} \Big(\frac{E^2}{\sin^2(zE)} - \frac{1}{z^2}\Big).$$

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#### Effective action asymptotics: Magnetic field

Consider the asymptotics of the effective action  $i\Gamma_{b^2}^{\text{eff}}(H, b)$ .

In the weak field,  $H \ll H_0$  (here  $H_0 = m^2/e = 4.41 \times 10^{13}$  Gs):

$$i\Gamma^{\mathrm{eff}}_{b_0^2}(H,b) \sim b_0^2 \frac{H^2}{m^2} + \mathcal{O}\left((H^2 b_0/m^3)^2\right), \quad \Gamma^{\mathrm{eff}}_{H-E}(H) \sim \frac{H^4}{m^4},$$

$$\frac{i\Gamma_{b_0^2}^{\text{eff}}(H,b)}{i\Gamma_{H-E}^{\text{eff}}(H)} \sim \left(\frac{b_0}{m}\right)^2 \left(\frac{H_0}{H}\right)^2.$$

In the strong field,  $H \gg H_0$ :

$$i\Gamma_{b_0^2}^{\text{eff}}(H,b) \sim b_0^2 H + \mathcal{O}\left(b_0^2 m^2 \log\left(\frac{H_0}{H}\right)\right), \quad i\Gamma_{H-E}^{\text{eff}}(H) \sim m^4 \left(\frac{H}{H_0}\right)^2 \log\left(\frac{H}{H_0}\right),$$
$$\frac{i\Gamma_{H-E}^{\text{eff}}(H)}{i\Gamma_{b_0^2}^{\text{eff}}(H,b)} \sim \frac{H}{H_0} \left(\frac{m}{b_0}\right)^2 \log\left(\frac{H}{H_0}\right).$$

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#### Effective action asymptotics: Magnetic field

Consider the asymptotics of the effective action  $i\Gamma_{b_{2}^{2}}^{\text{eff}}(H, b)$ .

In the weak field,  $H \ll H_0$  (here  $H_0 = m^2/e = 4.41 \times 10^{13}$  Gs):

$$i\Gamma^{\rm eff}_{b_0^2}(H,b) \sim b_0^2 {H^2 \over m^2} + \mathcal{O}\left((H^2 b_0/m^3)^2\right), \quad \Gamma^{\rm eff}_{H-E}(H) \sim {H^4 \over m^4},$$

$$\frac{i\Gamma_{b_0}^{\mathrm{eff}}(H,b)}{i\Gamma_{H-E}^{\mathrm{eff}}(H)} \sim \left(\frac{b_0}{m}\right)^2 \left(\frac{H_0}{H}\right)^2.$$

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$$\frac{i\Gamma_{H-E}^{\text{eff}}(H)}{i\Gamma_{b_0^2}^{\text{eff}}(H,b)} \sim \frac{H}{H_0} \left(\frac{m}{b_0}\right)^2 \log\left(\frac{H}{H_0}\right).$$

## Casimir effect within extended QED (O.G.Kharlanov and V.Ch.Zhukovsky)

- Vacuum energy via the zeta function regularization
- Vacuum energy via the residue theorem
- Conclusion

## 3 Effective action in QED under the Lorentz violation (A.F.Bubnov and V.Ch.Zhukovsky)

- The Model
- Induced Chern-Simons term in the constant field
- Quadratic contribution
- Results and conclusion

We have elaborated a new method of calculating the contribution of the CPTand Lorentz-violating correction  $b_{\mu}$  to the effective action of QED, which accounts for the external field exactly and is based on the proper time technique.

Using this method, we have obtained the following results:

- No effective Chern-Simons term linear in  $b^{\mu}$  in extended QED, in agreement with the publications:
  - Y.A. Sitenko, K.Y. Rulik ArXiv/hep-th/0212007,
  - B. Altschul ArXiv/hep-th/0602235,
- The  $b_0^2$ -term in the effective action, for the magnetic and the electric fields exactly taken into account.
- The Heisenberg-Euler-to- $b_0^2$  correction ratio  $\frac{i\Gamma_{H-E}^{\text{eff}}(H)}{i\Gamma_{eff}^{\text{eff}}(H,b)}$  is evaluated in

the weak- and strong-field cases.

References:

[2] V.Ch.Zhukovsky and A.Bubnov, to be published in arXiv/hep-th.

## Thank you for your attention!

