

*Sigma models for Lorentz group  
and  
superluminal propagation in 2d*

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work in progress

## *Is Lorentz symmetry exact?*

- Explore departures from LS
- Explicit violation of LS incompatible with general relativity
- Consider spontaneous LS breaking: the action is Lorentz invariant (and generally covariant), but the vacuum is not

# Sigma-model

$$SO(3, 1)/SO(3) \quad \rightarrow \quad V_\mu V^\mu = M^2$$

↑  
preserve spatial isotropy

## Einstein-aether

T. Jacobson, D. Mattingly (2001)


$$S = \int d^4x \left[ -\alpha_1 \partial_\mu V^\nu \partial^\mu V_\nu - \alpha_2 (\partial_\mu V^\mu)^2 \right. \\ \left. - \alpha_3 \epsilon^{\mu\nu\lambda\rho} \partial_\mu V_\nu \partial_\lambda V_\rho - \alpha_4 V^\mu \partial_\mu V^\nu V^\lambda \partial_\lambda V_\nu \right. \\ \left. + \lambda (V_\mu V^\mu - M^2) \right]$$


Requires UV completion:  $\Lambda \leq M$

2d toy-models may be useful: better quantum properties

C. Eling, T. Jacobson (2006)

$$S = \int d^2x \left( -\alpha_1 \partial_\mu V^\nu \partial^\mu V_\nu - \alpha_2 \partial_\mu V^\mu \partial_\nu V^\nu \right. \\ \left. - \alpha_3 \partial_\mu V^\mu \epsilon^{\nu\lambda} \partial_\nu V_\lambda + \lambda (V^\mu V_\mu - 1) \right)$$

violates parity 

$V_\mu V^\mu = 1$  

Work in light-cone coordinates  $x^\pm = \frac{1}{\sqrt{2}}(t \pm x)$

Introduce “rapidity” field:  $V^\pm = \frac{1}{\sqrt{2}}e^{\pm\psi}$

$$S = \int dx^+ dx^- \frac{1}{g^2} \left\{ \partial_+ \psi \partial_- \psi + \frac{\beta_{(+)}}{2} (\partial_+ \psi)^2 e^{2\psi} + \frac{\beta_{(-)}}{2} (\partial_- \psi)^2 e^{-2\psi} \right\}$$

$g^2 = \frac{1}{2\alpha_1 + \alpha_2}$

$\beta_{(\pm)} = \frac{-\alpha_2 \pm \alpha_3}{2\alpha_1 + \alpha_2}$

- Lorentz symmetry is realized non-linearly:

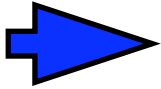
$$\psi(x^+, x^-) \mapsto \psi(e^\gamma x^+, e^{-\gamma} x^-) + \gamma$$

- Renormalizable by power-counting

# One-loop RG flow

$g^2$  does not run

$$\frac{d\beta_{(\pm)}}{d \log \Lambda} = - \frac{g^2 \beta_{(\pm)}}{\pi \sqrt{1 - \beta_{(+)}\beta_{(-)}}$$

- In UV:  $\beta_{(\pm)} \rightarrow 0$   theory flows to the **free Lorentz invariant** limit
- In IR: three cases

$$\beta_{(+), \beta_{(-)} : ++, +0, +-$$

# One-loop RG flow

$$\beta_{(+)} = \beta_{(-)} = \beta$$

$$\frac{d\beta}{d \log \Lambda} = - \frac{g^2 \beta}{\pi \sqrt{1 - \beta^2}}$$

Infrared pole  $\beta = 1$  at finite scale  strong coupling in IR

$$\beta_{(-)} = 0$$

No physical running: change of  $\beta$  is compensated by the shift of  $\psi$


# One-loop running

$$\beta_{(+)} = -\beta_{(-)} = \beta$$

$$\frac{d\beta}{d \log \Lambda} = -\frac{g^2 \beta}{\pi \sqrt{1 + \beta^2}}$$

In IR  $\beta \rightarrow \infty$

$$S = \frac{1}{2\kappa^2} \int dx_+ dx_- \left[ (\partial_+ \psi)^2 e^{2\psi} - (\partial_- \psi)^2 e^{-2\psi} \right]$$

$$\kappa^2 = g^2 / \beta$$


$$\frac{d\kappa}{d \log \Lambda} = \frac{\kappa^3}{2\pi}$$

theory flows to a weakly coupled point

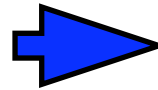


# Propagation of signals

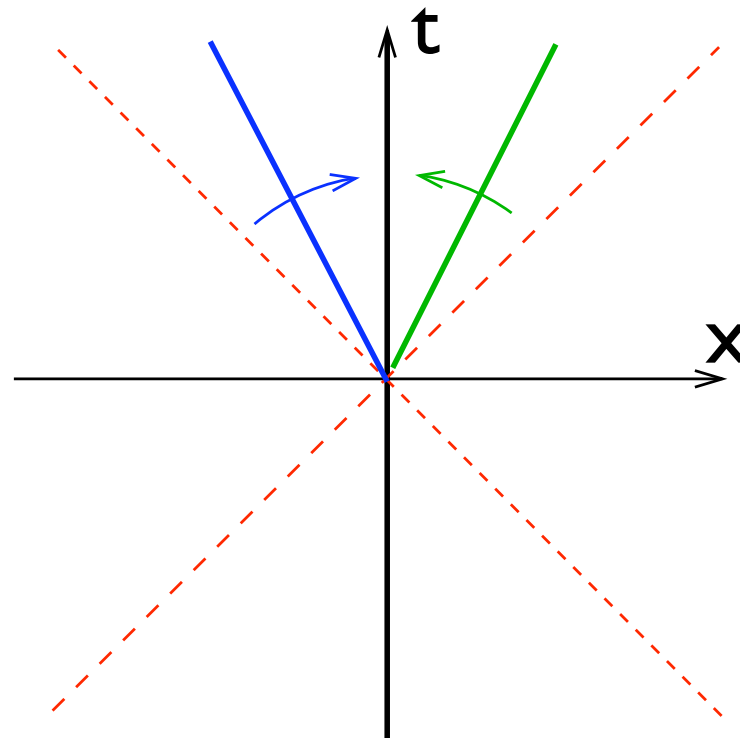
$$\beta_{(+)} = \beta_{(-)} = \beta$$

Fix background

$$\psi = 0$$



$$v^2 = \frac{1 - \beta}{1 + \beta}$$

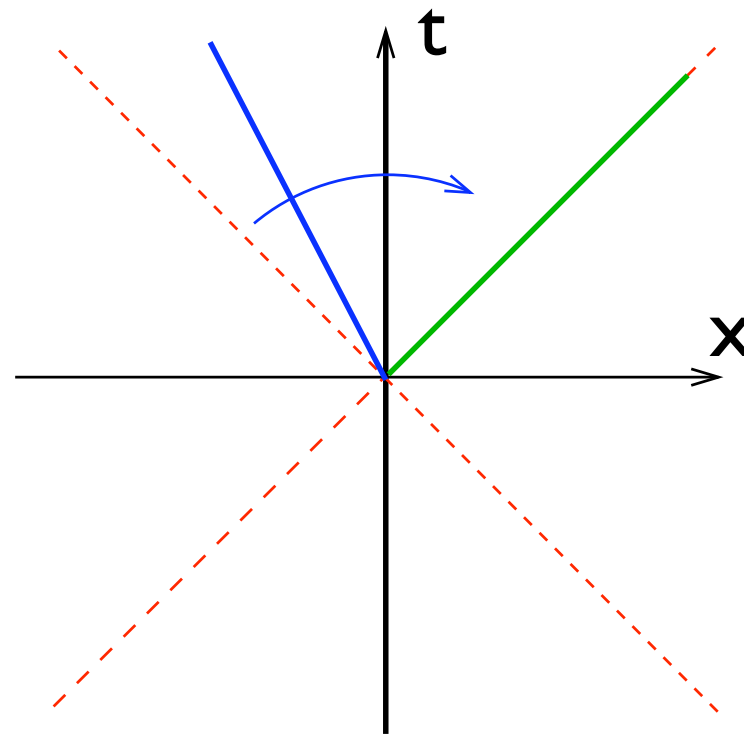


# Propagation of signals

$$\beta_{(-)} = 0$$

$$v_L = \frac{\beta - 2}{\beta + 2}$$

$$v_R = 1$$

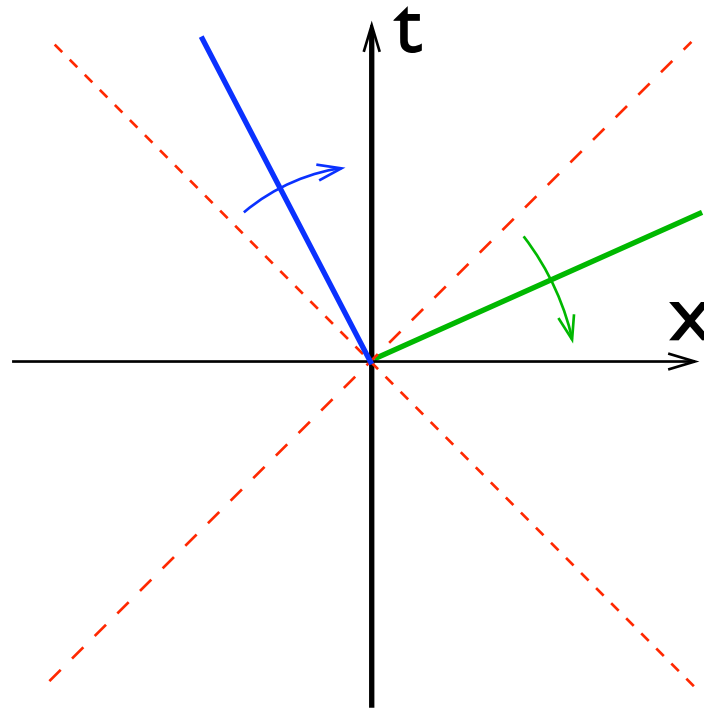


# Propagation of signals

$$\beta_{(+)} = -\beta_{(-)} = \beta$$

$$v_L = -\beta + \sqrt{1 + \beta^2}$$

$$v_R = \beta + \sqrt{1 + \beta^2}$$



The right-moving mode is **superluminal**

# Restoration of Lorentz invariance

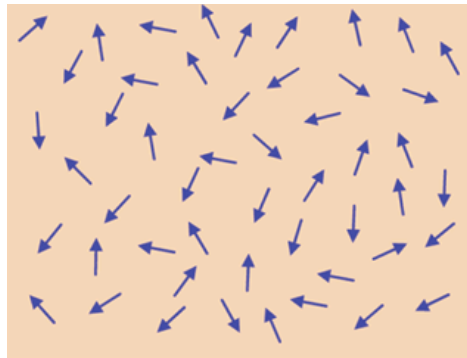
No spontaneous breaking of continuous symmetries in 2d

N.D. Mermin, H. Wagner (1966)

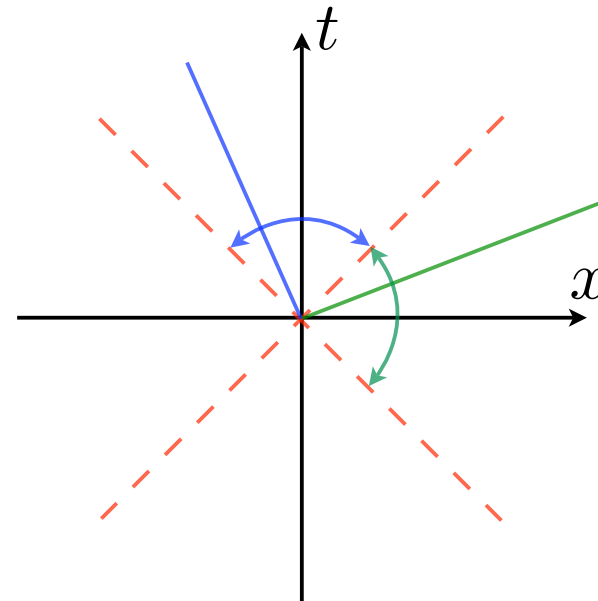
S. Coleman (1973)

Average over classical vacua.

Example: 2d ferromagnet

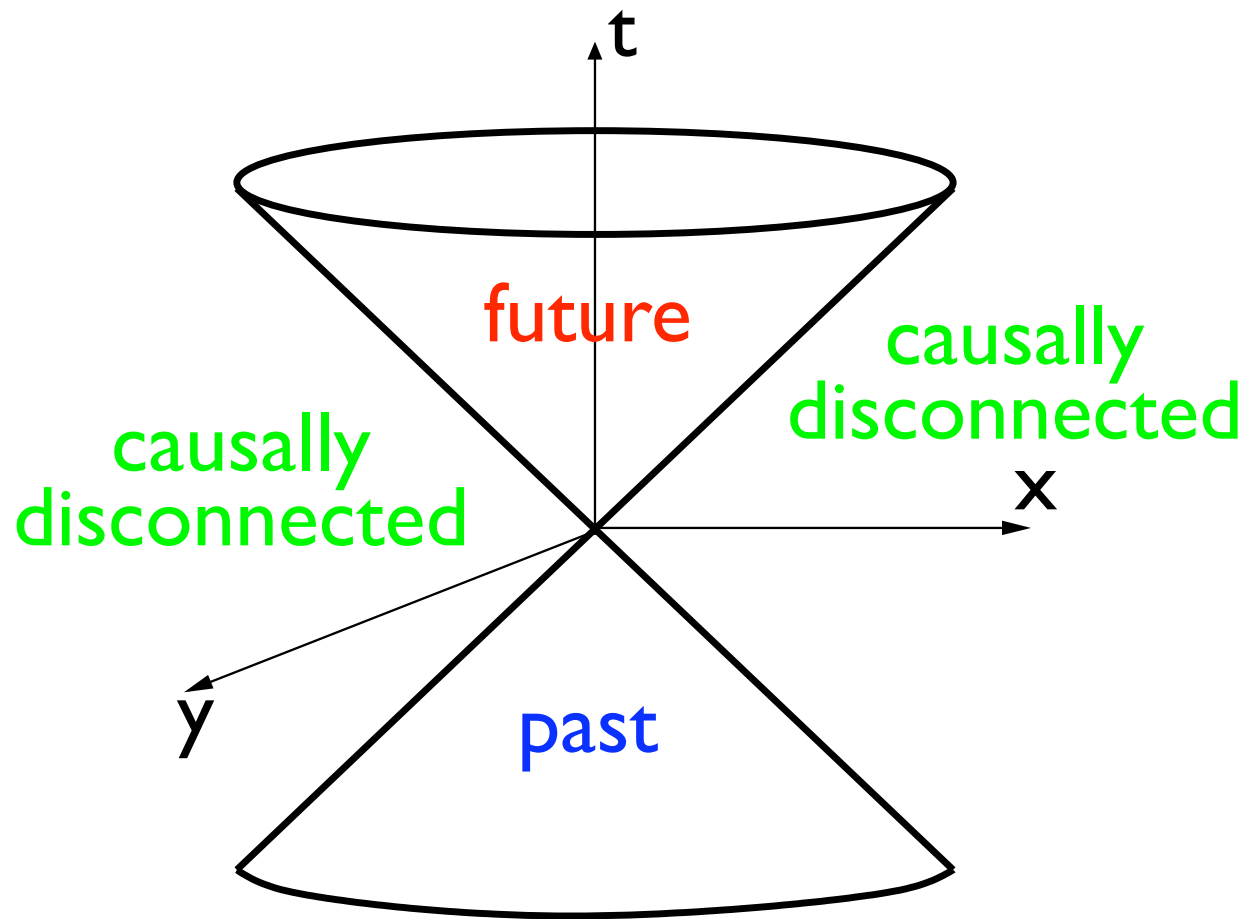


Superluminal Einstein-aether

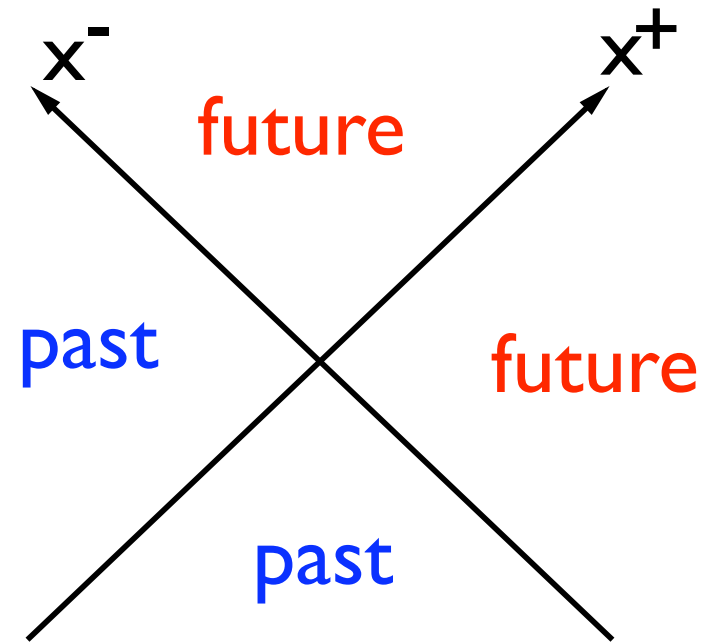
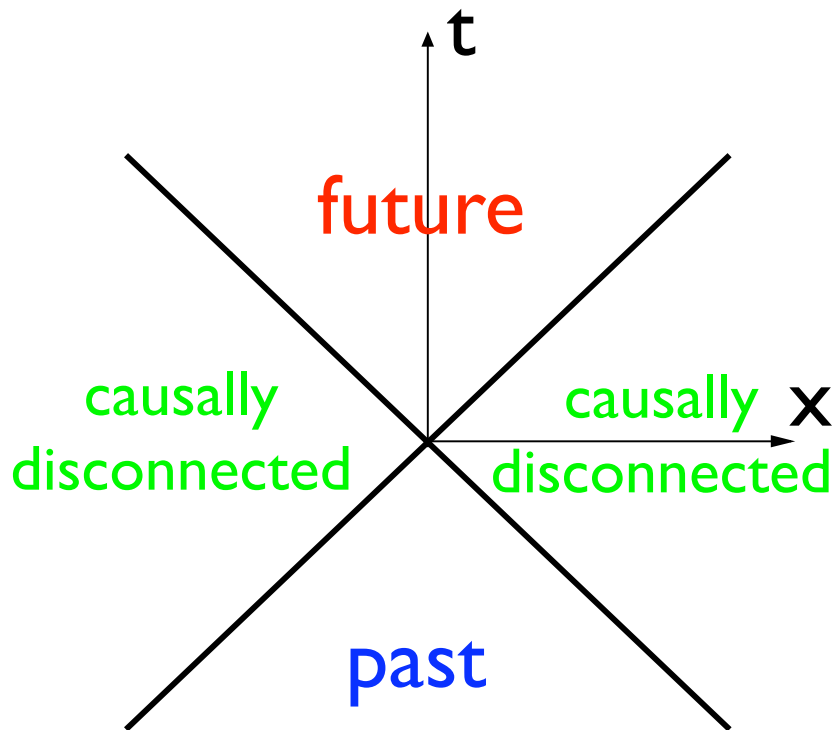


Instead of spontaneous LS breaking a *Lorentz invariant* model with *superluminal propagation*

# Lorentz invariant causal structure: $d > 2$



# Lorentz invariant causal structures: 2d



Positive-definite Hamiltonian for evolution along  $x^+$

$$H_{x^+} = \int dx^- \frac{\beta}{2g^2} \left[ (\partial_+ \psi)^2 e^{2\psi} + (\partial_- \psi)^2 e^{-2\psi} \right]$$

## Hints at integrability: $\beta_{(-)} = 0$

$$S = \int dx^+ dx^- \frac{1}{g^2} \left[ \partial_+ \psi \partial_- \psi + \frac{\beta}{2} (\partial_+ \psi)^2 e^{2\psi} \right]$$

- symmetry:

$$\psi(x^+, x^-) \mapsto \psi(f(x^+), g(x^-)) - \frac{1}{2} \log f'(x^+) + \frac{1}{2} \log g'(x^-)$$

$$f = \frac{ax^+ + b}{cx^+ + d}$$

arbitrary

**N.B. half** of conformal symmetry

- Infinite number of integrals of motion  $\partial_+ q = 0$

$$q = (\partial_- \psi)^2 - \partial_-^2 \psi - \frac{\beta}{2} \partial_- \partial_+ \psi e^{2\psi}$$

- General classical solution
- Renormalizable by normal ordering

# Hints at integrability: general case

Equation of motion:

$$2\partial_+\partial_-\psi + \beta_{(+)}(\partial_+^2\psi + (\partial_+\psi)^2)e^{2\psi} + \beta_{(-)}(\partial_-^2\psi - (\partial_-\psi)^2)e^{-2\psi} = 0$$

equivalent to zero-curvature condition:

$$[\partial_+ + V, \partial_- + U] = 0$$

$$U = \sigma_+ + \frac{1}{\lambda_- - \lambda_+} e^{-a\sigma_-} \left( \frac{\lambda_+^2 \sigma_+}{\lambda_+ - \lambda_-} \right) e^{a\sigma_-} + \frac{1}{\lambda_- - \lambda_-} e^{-b\sigma_-} \left( \frac{-\lambda_-^2 \sigma_+}{\lambda_+ - \lambda_-} \right) e^{b\sigma_-}$$

$$V = \frac{1}{\lambda_- - \lambda_+} e^{-a\sigma_-} \left( \frac{\lambda_+ s \sigma_+}{\lambda_+ - \lambda_-} \right) e^{a\sigma_-} + \frac{1}{\lambda_- - \lambda_-} e^{-b\sigma_-} \left( \frac{-\lambda_- s \sigma_+}{\lambda_+ - \lambda_-} \right) e^{b\sigma_-}$$

$$a = -\partial_-\psi + \lambda_- \partial_+\psi e^{2\psi} \quad b = -\partial_-\psi + \lambda_+ \partial_+\psi e^{2\psi} \quad s = e^{-2\psi}$$

$$\beta_{(+)} = \frac{-2\lambda_+\lambda_-}{\lambda_+ + \lambda_-} \quad \beta_{(-)} = \frac{-2}{\lambda_+ + \lambda_-} \quad \sigma_+ = \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix}$$

$$\sigma_- = \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$$



# Coupling to gravity

Start from the Einstein-aether form of the action

$$S_{gr} = -\frac{1}{2\pi\kappa} \int d^2x \sqrt{-g} \left( R + \mu^2 \right) + S_{EA}(g_{\mu\nu}, V_\mu)$$

Fix conformal gauge  Einstein-aether sector contributes to the Liouville action as a single scalar boson

+ explicit coupling to the Liouville field

$$S = \tilde{S}_{EA}(\phi, \psi) + S_L(\phi) + \dots$$

$$\tilde{S}_{EA} = \int d^2x \left\{ \frac{1}{g^2} \partial_+ \psi \partial_- (\psi - \phi) + \frac{\beta_{(+)}}{2g^2} e^{2\psi - \phi} (\partial_+ \psi)^2 + \frac{\beta_{(-)}}{2g^2} e^{\phi - 2\psi} (\partial_- (\psi - \phi))^2 \right\}$$

# Prospects

- Toy models for issues related to causality

Weak coupling to ordinary (massive) fields

$$S_\chi = \int d^2x \left( \partial_+ \chi \partial_- \chi - \frac{m^2 \chi^2}{2} + \frac{\gamma(+)}{2} (\partial_+ \chi)^2 e^{2\psi} + \frac{\gamma(-)}{2} (\partial_- \chi)^2 e^{-2\psi} \right)$$

➡ apparent “a-causal effects”

- Toy models for extraction of information from black holes

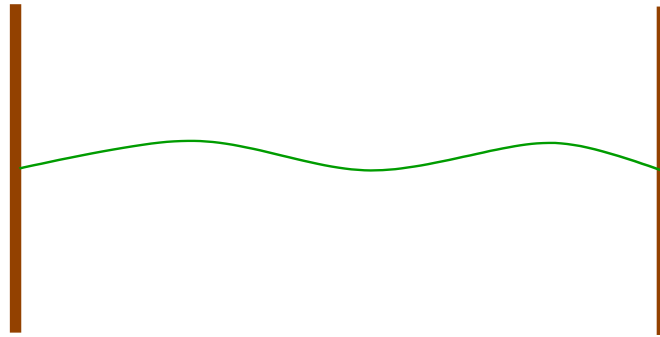
## Prospects: more ambitious

Consider EA as (a part of) world-sheet action of a string

Possible action:

$$S = \int d^2x \left\{ \frac{1}{g^2} [\partial_+ \psi \partial_- (\psi - \phi) - \partial_+ X^i \partial_- X^i] \right. \\ \left. + \frac{\beta}{2g^2} e^{2\psi - \phi} [(\partial_+ \psi)^2 + (\partial_+ X^i)^2] \right\} + S_L(\phi) + \dots$$

The resulting theory will be UV complete, contain gravity and possess unusual properties: **a-causality ? non-locality ?**



If it exists ...

At stake: holidays at  $\alpha$  Centauri

