



# Exceptional Symmetries in the Light- Cone Gauge and Possible Counter Terms.

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When we study supersymmetric theories,  
 $N = 4$  Yang-Mills or  $N = 8$  Supergravity or  
their higher-dimensional versions always pop  
up.

In the light-cone formulation  
their superfields are particularly simple.

One can regard them as master fields for a  
series of field theories.

In light-cone gauge field theory they can be  
treated similarly.

## Light-Frame Formulation

Dirac showed that any direction within the light-cone can be "time".

Choose  $x^+ = \frac{1}{\sqrt{2}}(x^0 + x^3)$  as the time.

The coordinates and the derivatives that we will use will then be

$$\begin{aligned}x^\pm &= \frac{1}{\sqrt{2}}(x^0 \pm x^3) \\ \partial^\pm &= \frac{1}{\sqrt{2}}(-\partial_0 \pm \partial_3) \\ x &= \frac{1}{\sqrt{2}}(x_1 + i x_2) \\ \bar{\partial} &= \frac{1}{\sqrt{2}}(\partial_1 - i \partial_2) \\ \bar{x} &= \frac{1}{\sqrt{2}}(x_1 - i x_2) \\ \partial &= \frac{1}{\sqrt{2}}(\partial_1 + i \partial_2)\end{aligned}$$

We will only consider massless theories so we solve the condition  $p^2 = 0$ . We then find

$$p^- = \frac{p\bar{p}}{p^+}.$$

The generator  $p^-$  is really the **Hamiltonian**.

In the light-cone frame the supersymmetry generator  $Q$  splits up into two two-component spinors that can be rewritten as two complex operators, which we call

$$Q_+ = -\frac{1}{2}\gamma_+\gamma_-Q \text{ and } Q_- = -\frac{1}{2}\gamma_-\gamma_+Q.$$

The light-cone supersymmetry algebra is then

$$\begin{aligned}\{Q_+^m, \bar{Q}_{+n}\} &= -\sqrt{2}\delta_n^m P^+ \\ \{Q_-^m, \bar{Q}_{-n}\} &= -\sqrt{2}\delta_n^m P^- \\ \{Q_+^m, \bar{Q}_{-n}\} &= -\sqrt{2}\delta_n^m P,\end{aligned}$$

The superPoincaré algebra can now be represented on a superspace with coordinates

$$x^\pm, \quad x, \quad \bar{x}, \quad \theta^m, \quad \bar{\theta}_n$$

**All generators with a minus-component get non-linear contributions**

We will denote the derivatives of the  $\theta$ 's

$$\bar{\partial}_m \equiv \frac{\partial}{\partial \theta^m} ; \quad \partial^m \equiv \frac{\partial}{\partial \bar{\theta}_m}$$

The kinematical  $q$ 's will be represented by

$$q_+^m = -\partial^m + \frac{i}{\sqrt{2}} \theta^m \partial^+ , \quad \bar{q}_{+n} = \bar{\partial}_n - \frac{i}{\sqrt{2}} \bar{\theta}_n \partial^+ ,$$

and the dynamical ones as

$$q_-^m = \frac{\bar{\partial}}{\partial^+} q_+^m , \quad \bar{q}_{-m} = \frac{\partial}{\partial^+} \bar{q}_{+m} .$$

On this space we can also represent "chiral" derivatives anticommuting with the supercharges  $Q$ .

$$d^m = -\partial^m - \frac{i}{\sqrt{2}} \theta^m \partial^+ , \quad \bar{d}_n = \bar{\partial}_n + \frac{i}{\sqrt{2}} \bar{\theta}_n \partial^+ .$$

To find an irreducible representation we have to impose the **chiral constraints**

$$d^m \phi = 0 ; \quad \bar{d}_m \bar{\phi} = 0 ,$$

on a complex superfield  $\phi(x^\pm, x, \bar{x}, \theta^m, \bar{\theta}_n)$ . The solution is then that

$$\phi = \phi(x^+, y^- = x^- - \frac{i}{\sqrt{2}} \theta^m \bar{\theta}_m, x, \bar{x}, \theta^m).$$

It is particularly interesting to study the cases  $N = 4 \times \text{integer}$ . For those values one can impose a further condition on the superfield  $\phi$  namely the "**inside out**" condition

$$\bar{d}_{m_1} \bar{d}_{m_2} \dots \bar{d}_{m_{N/2-1}} \bar{d}_{m_{N/2}} \phi = \frac{1}{2} \epsilon_{m_1 m_2 \dots m_{N-1} m_N} d^{m_{N/2+1}} d^{m_{N/2+2}} \dots d^{m_{N-1}} d^{m_N} \bar{\phi}$$

When  $\frac{N}{4}$  is odd, the superfield has to transform as the adjoint representation of an external group with structure constants  $f^{abc}$ .

$N = 4$

$$\begin{aligned} \phi(y) &= \frac{1}{\partial^+} A(y) + \frac{i}{\sqrt{2}} \theta^m \theta^n \bar{C}_{mn}(y) \\ &+ \frac{1}{12} \theta^m \theta^n \theta^p \theta^q \epsilon_{mnpq} \partial^+ \bar{A}(y) \\ &+ \frac{i}{\partial^+} \theta^m \bar{\chi}_m(y) + \frac{\sqrt{2}}{6} \theta^m \theta^n \theta^p \epsilon_{mnpq} \chi^q(y) . \end{aligned}$$

$N = 8$

$$\begin{aligned} \phi(y) &= \frac{1}{\partial^{+2}} h(y) + i \theta^m \frac{1}{\partial^{+2}} \bar{\chi}_m(y) \\ \dots &+ \theta^{mnp r} \bar{C}_{mnp r}(y) \\ \dots &+ \tilde{\theta}_m^{(7)} \partial^+ \chi^m(y) + \tilde{\theta}^{(8)} \partial^{+2} \bar{h}(y) , \end{aligned}$$



## The $N = 4$ Yang-Mills Theory

This was the first action we constructed

$$\begin{aligned} \mathcal{S} = & - \int d^4x \int d^4\theta d^4\bar{\theta} \\ & \left\{ \bar{\phi}^a \frac{\square}{\partial^+{}^2} \phi^a + \frac{4g}{3} f^{abc} \left( \frac{1}{\partial^+} \bar{\phi}^a \phi^b \bar{\partial} \phi^c + \text{c.c.} \right) \right. \\ & - g^2 f^{abc} f^{ade} \left( \frac{1}{\partial^+} (\phi^b \partial^+ \phi^c) \frac{1}{\partial^+} (\bar{\phi}^d \partial^+ \bar{\phi}^e) \right. \\ & \left. \left. + \frac{1}{2} \phi^b \bar{\phi}^c \phi^d \bar{\phi}^e \right) \right\}. \end{aligned}$$

With this action we (Brink, Lindgren and Nilsson 1982) proved that the [perturbation expansion is finite](#).

The  $N = 8$  Supergravity action to first order is then

$$\int d^4x \int d^8\theta d^8\bar{\theta} \mathcal{L} \equiv \int \mathcal{L},$$

where,

$$\mathcal{L} = -\bar{\phi} \frac{\square}{\partial^{+4}} \phi + \left( \frac{4\kappa}{3\partial^{+4}} \bar{\phi} \bar{\partial} \bar{\partial} \phi \partial^{+2} \phi - \frac{4\kappa}{3\partial^{+4}} \bar{\phi} \bar{\partial} \partial^+ \phi \bar{\partial} \partial^+ \phi + c.c. \right)$$

How do we construct the **four-point function**?

We can do it by trial and error.

Too hard.

Instead we found a remarkable property of maximally supersymmetric theories.

(with Ananth and Ramond)

## The Hamiltonian as a Quadratic Form

The usual relation is that

$$H = \frac{1}{4}\{Q_-^m, Q_{-m}\}$$

For both  $N = 4$  and  $N = 8$

$$H = \int \delta_{\bar{q}_{-m}} \bar{\phi} \delta_{q_{-m}} \phi$$

Not an anticommutator, but a quadratic form.

With this form we could run a Mathematica program comparing with the four-point function of gravity.

The result was a four-point coupling with **96** terms. (In the covariant form there are about **5000** terms.)  
with Ananth, Heise and Svendsen.

## Higher Symmetries for $N = 4$ Yang-Mills Theory

We know that the  $d = 4$  theory is **conformally invariant**, i.e. under  $PSU(2,2|4)$  even for the quantum case. We can in fact construct the whole theory by closing the **conformal algebra** by guessing the correct dynamical supersymmetry generator  $Q_-$ .

**This scheme can be followed for all superconformally invariant field theories, also for  $d=3$  or  $d=6$ . I had planned to talk about  $d=3$ , but we still miss something there.**

(With Kim and Ramond)

## Higher Symmetries for $N = 8$ Supergravity Theory

$N = 8$  Supergravity, unlike  $N = 4$  Yang-Mills, is not superconformal invariant; however, it does have the non-linear Cremmer-Julia  $E_{7(7)}$  symmetry.

### How do we implement the $E_{7(7)}$ symmetry?

Go back to covariant component form (Cremmer, Julia and Freedman, de Wit)

$$\mathcal{L} = \mathcal{L}_S + \mathcal{L}_V + \mathcal{L}_{\text{others}}$$

$\mathcal{L}_S$  is a Coleman-Wess-Zumino non-linear Lagrangian. The  $E_{7(7)}$  is clear.

$\mathcal{L}_V$  can be written as

$$\mathcal{L}_V = -\frac{1}{8} F^{\mu\nu ij} G_{\mu\nu}^{ij} ,$$

The Lagrangian is quadratic in the field strengths.  
Introduce the *self-dual complex* field strengths

$$\mathbb{F}^{\mu\nu ij} = \frac{1}{2} (F^{\mu\nu ij} + i\tilde{F}^{\mu\nu ij})$$

and

$$\mathbb{G}^{\mu\nu ij} = \frac{1}{2} (G^{\mu\nu ij} + i\tilde{G}^{\mu\nu ij})$$

The equations of motion are given by

$$\partial_\mu G^{\mu\nu ij} = \partial_\mu (\mathbb{G}^{\mu\nu ij} + \bar{\mathbb{G}}^{\mu\nu ij}) = 0 ,$$

while the Bianchi identities read

$$\partial_\mu \tilde{F}^{\mu\nu ij} = \partial_\mu (\mathbb{F}^{\mu\nu ij} - \bar{\mathbb{F}}^{\mu\nu ij}) = 0 .$$

Assemble in one column vector with **56** complex entries

$$Z^{\mu\nu} = \begin{pmatrix} \mathbb{G}^{\mu\nu ij} + \mathbb{F}^{\mu\nu ij} \\ \mathbb{G}^{\mu\nu ij} - \mathbb{F}^{\mu\nu ij} \end{pmatrix} \equiv \begin{pmatrix} X^{\mu\nu ab} \\ Y^{\mu\nu}_{ab} \end{pmatrix},$$

where  $a, b$  are  $SU(8)$  indices, with upper(lower) antisymmetric indices for **28(28)**.

This is a **56** under  $E_{7(7)}$ .

The 70 transformations are

$$\begin{aligned} \delta X^{\mu\nu ab} &= \Xi^{abcd} Y^{\mu\nu}_{cd} \\ \delta Y^{\mu\nu}_{ab} &= \Xi_{abcd} X^{\mu\nu cd}, \end{aligned}$$

We now specialize to the **light-cone gauge**. We choose  $A^+ = 0$  and solve for  $A^-$ . We then make non-linear **field redefinitions**,  $A^{ij} \rightarrow B^{ij}$  and  $C^{ijkl} \rightarrow D^{ijkl}$  to get rid of "time" derivatives" in the interaction terms.

This will mix up the fields and the Hamiltonian is no longer quadratic in  $B^{ij}$ .

We can now read off the  $E_{7(7)}/SU(8)$  transformations in the vector and scalar fields.

**However, the other fields now take part in the transformations!**



The  $\frac{E_{7(7)}}{SU(8)}$  quotient symmetry must commute with the other symmetries in particular with the supersymmetry.  $[\delta_{70}, \delta_S]\varphi = 0$ .

(There is no  $E_{7(7)}$  supergroup.)

By using that we get the transformations for all fields in the multiplet.

How can  $\frac{E_{7(7)}}{SU(8)}$  commute when  $SU(8)$  does not, and

$$[\delta_{70}, \delta_{70}] = \delta_{SU(8)}?$$

Consider the Jacobi identity

$$([\delta_{70}, \delta_{70}], \delta_S) + ([\delta_S, \delta_{70}], \delta_{70}) + ([\delta_{70}, \delta_S], \delta_{70})\varphi = 0$$

Since  $[\delta_S, \delta_{70}]\delta_{70}\varphi \neq 0$ , it works!  $\delta_{70}\varphi$  non-linear! We only claim that  $[\delta_S, \delta_{70}]\varphi = 0$ . All fields including the graviton transform under  $\frac{E_{7(7)}}{SU(8)}$  and into each other.

## Some of the transformations

Vectors:

$$\begin{aligned}
 \delta \bar{B}_{ij} = & -\kappa \Xi^{klmn} \left( \frac{1}{4} \bar{D}_{ijkl} \bar{B}_{mn} + \frac{1}{4!} \frac{1}{\partial^+} \bar{D}_{klmn} \partial^+ \bar{B}_{ij} - \frac{1}{4!} \epsilon_{ijklmnr} \frac{1}{\partial^+} B^{rs} \partial^+ h \right. \\
 & \left. + \frac{i}{3!} \frac{1}{\partial^+} \bar{\chi}_{klm} \bar{\chi}_{ijn} - \frac{i}{3!} \epsilon_{ijklmnr} \frac{1}{\partial^+} \chi^{rst} \bar{\psi}_n \right) \\
 & + \kappa \bar{\Xi}_{ijkl} \frac{1}{\partial^+} \left( \frac{1}{4} D^{klmn} \partial^+ \bar{B}_{mn} - \frac{1}{\partial^+} B^{kl} \partial^{+2} h \right. \\
 & \left. + \frac{i}{4(3!)^2} \bar{\chi}_{mnp} \bar{\chi}_{rst} \epsilon^{klmnpqrst} - 3i \frac{1}{\partial^+} \chi^{kln} \partial^+ \bar{\psi}_n \right) \quad (1)
 \end{aligned}$$

Gravitini:

$$\begin{aligned}
 \delta \bar{\psi}_i = & -\kappa \bar{\Xi}^{mnpq} \left( \frac{1}{4! \cdot 3!} \epsilon_{mnpqirst} D^{rstu} \bar{\psi}_u + \frac{1}{4!} \frac{1}{\partial^+} \bar{D}_{mnpq} \partial^+ \bar{\psi}_i + \frac{1}{4!} \bar{D}_{mnpq} \bar{\psi}_i \right. \\
 & \left. - \frac{1}{4!} \epsilon_{mnpqirst} \frac{1}{\partial^+} \chi^{rst} \partial^+ h + \frac{1}{4} \bar{\chi}_{imn} \bar{B}_{pq} + \frac{1}{3!} \frac{1}{\partial^+} \bar{\chi}_{mnp} \partial^+ \bar{B}_{iq} \right) \quad (2)
 \end{aligned}$$

Gravition:

$$\delta h = -\kappa \Xi^{mnpq} \left( \frac{1}{4!} \frac{1}{\partial^+} \bar{D}_{mnpq} \partial^+ h + \frac{1}{8} \bar{B}_{mn} \bar{B}_{pq} + \frac{i}{\partial^+} \bar{\chi}_{mnp} \bar{\psi}_q \right) \cdot \quad (3)$$

We then find that we can write the order  $\kappa$  transformation as

$$\delta\varphi = \frac{\kappa}{4!} \equiv^{mnpq} \frac{1}{\partial+2} (\bar{d}_m \bar{d}_n \bar{d}_p \bar{d}_q \frac{1}{\partial+} \varphi \partial^{+3} \varphi - 4 \bar{d}_m \bar{d}_n \bar{d}_p \varphi \bar{d}_q \partial^{+2} \varphi + 3 \bar{d}_m \bar{d}_n \partial^{+} \varphi \bar{d}_p \bar{d}_q \partial^{+} \varphi) + \dots .$$

This expression is in fact unique! It can be rewritten in a very efficient form

$$\frac{\kappa}{4!} \equiv^{mnpq} \left( \frac{\partial}{\partial\eta} \right)_{mnpq} \frac{1}{\partial+2} \left( e^{\eta \hat{d}} \partial^{+3} \varphi e^{-\eta \hat{d}} \partial^{+3} \varphi \right) \Big|_{\eta=0} ,$$

where  $\hat{d} = \frac{\bar{d}}{\partial+}$ .

## The Hamiltonian

We write

$$\delta_s^{dyn} \varphi = \delta_s^{dyn(0)} \varphi + \delta_s^{dyn(1)} \varphi + \delta_s^{dyn(2)} \varphi + \mathcal{O}(\kappa^3)$$

We can now require

$$[\delta_{70}, \delta_s^{dyn}] \varphi = 0$$

Here we can use the inhomogeneity of the 70 transformation

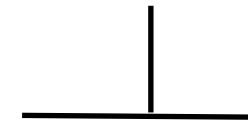
$$[\delta_{70}^{(-1)}, \delta_s^{dyn(2)}] \varphi + [\delta_{70}^{(1)}, \delta_s^{dyn(0)}] \varphi = 0$$

This gives the order  $\kappa^2$  dynamical supersymmetry. We can use the quadratic form to find the *Hamiltonian* to order  $\kappa^2$ . Much simpler than before!

## Possible counterterms for N=8

Let us check first in gravity. We can write the three point coupling as

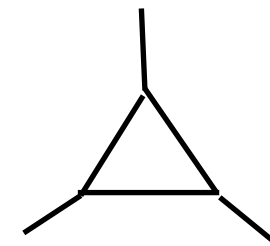
$$\begin{aligned}\delta_H^\kappa h &= \kappa \partial^{+n} \left[ e^{a\hat{\partial}} \partial^{+m} h e^{-a\hat{\partial}} \partial^{+m} h \right] \Big|_{a^2} \\ &\equiv \kappa \partial^{+n} \left( \frac{\partial}{\partial a} \right)^2 \left[ e^{a\hat{\partial}} \partial^{+m} h e^{-a\hat{\partial}} \partial^{+m} h \right] \Big|_{a=0},\end{aligned}$$



A possible one-loop counter term is

$$\delta_H^{g_1} h = \kappa^3 \partial^{+n} \left[ E \partial^{+m} h E^{-1} \partial^{+m} h \right] \Big|_{a^3, b},$$

$$E = e^{a\hat{\partial} + b\hat{\partial}} \quad \text{and} \quad E^{-1} = e^{-a\hat{\partial} - b\hat{\partial}},$$



Consistent with the algebra for two choices of  $m$  and  $n$

This can in fact be generalized to all orders.

$$\delta_H^{g_l} h = \kappa^{2l+1} \partial^+ [E \partial^{+l} h E^{-1} \partial^{+l} h] \Big|_{a^{2+l}, b^l},$$

$$\delta_H^{g_l} h = \kappa^{2l+1} \frac{1}{\partial^{+3}} [E \partial^{+(l+2)} h E^{-1} \partial^{+(l+2)} h] \Big|_{a^{2+l}, b^l}.$$

There is another series starting with

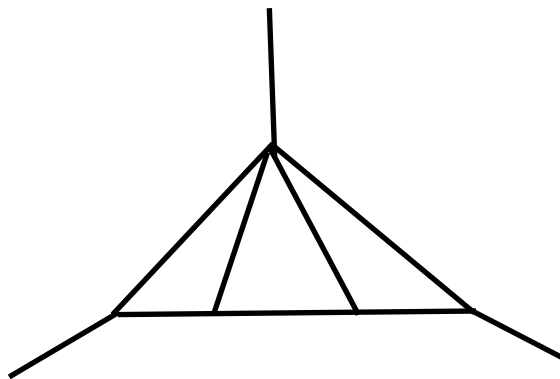
$$\delta_H^{g_2} h = \kappa^5 \frac{1}{\partial^{+3}} [E \partial^{+4} \bar{h} E^{-1} \partial^{+4} \bar{h}] \Big|_{b^6}.$$

We are interested in counterterms which are non-zero when we use the equation of motion.

$$\partial^- h = \delta_H h = \frac{\partial \bar{\partial}}{\partial^+} h + \mathcal{O}(\kappa).$$

All but the third terms can be written as  $\square(..h...h)$

We only find a two-loop three-point counter term.  
Consider a three-loop term



Goes like

$$\int dp^{12} \frac{p^{10}}{p^{14}} \sim p^8$$

Terribly divergent but must be  $\sim p_i^2$  where  
 $i = 1, 2, \text{ or } 3$ .

## N=8 Supergravity

There are no three-point counter terms for  $N = 8$

$$\delta_H \phi = \dots(\dots \bar{\phi} \bar{\phi})$$

since the r.h.s. is not chiral.

When we consider the four-point coupling we have to use the  $E_{7(7)}$  symmetry. Remember how we obtain the four-point coupling.

$$[\delta_{70}^{(-1)}, \delta_s^{dyn(2)}] \varphi + [\delta_{70}^{(1)}, \delta_s^{dyn(0)}] \varphi = 0$$

The terms talk to each other pairwise. They have the same number of derivatives.



A four-point counterterm  $\delta_{s,c}^{dyn(2)}$  must satisfy

$$[\delta_{70}^{(-1)}, \delta_{s,c}^{dyn(2)}] \varphi = 0$$

Furthermore it has to satisfy all the commutations rules with the full  $N = 8$  superalgebra.

Well-defined problem but algebraically difficult.

We still do not have the final result.